

Exam

Hour Exam 2: Wednesday, October 25th

- In-class, covering waves, electromagnetism, and relativity
- Twenty multiple-choice questions
- Will cover: Chapters 8, 9 10 and 11
Lecture material
- You should bring
 - 1 page notes, written single sided
 - #2 Pencil and a Calculator
 - Review Monday October 23rd
 - Review test will be available online on Monday

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From before

$$E = \gamma m_0 c^2 = KE_{rel} + m_0 c^2$$

$$E = \gamma m_0 c^2, \text{ or } E = mc^2$$

- The total energy of a particle is dependent on it's kinetic energy and its mass.
- Even when the particle is not moving it has energy.
- **Mass is another form of energy**
 - Moreover, energy can show up as mass.
 - The energy to put all the protons together in a nucleus gives the nucleus more mass!

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Space/Time - Energy/Momentum

- Relativity mixes up space and time - also energy and momentum
 - When converting from one inertial frame to another
 - In the time dilation and length contraction formulas time is in the length formula and length is in the time volume through the velocity (length/time)
 - In the total energy formula momentum(or kinetic energy) and mass energy are related
- There are combinations of space/time and energy/momentum that observers in any inertial frame will measure the as the same
 - For energy and momentum this invariant says that all observers can agree on mass an object has when it's at rest!

$$x^2 - c^2 t^2$$

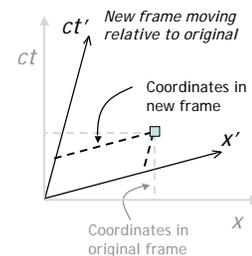
$$E^2 - c^2 p^2 = (m_0 c^2)^2$$

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Observing from a new frame

- In relativity these events will look different in reference frame moving at some velocity
- The new reference frame can be represented as same events along different coordinate axes
- **A graphical way of showing that length and time are contracted or expanded.**



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The Equivalence Principle



Clip from
*Einstein Nova
special*

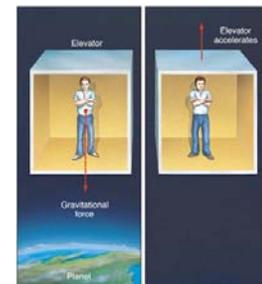
- Led Einstein to postulate the Equivalence Principle

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Equivalence principle

Accelerating reference frames are **indistinguishable** from a gravitational force



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Try some experiments

Constant velocity

Constant accel.

$t=0$ $t=t_0$ $t=2t_0$ $t=0$ $t=t_0$ $t=2t_0$

Floor accelerates upward to meet ball

Cannot do any experiment to distinguish accelerating frame from gravitational field

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Light follows the same path

Path of light beam in our frame

Velocity = v

Velocity = $v+at_0$

Velocity = $v+2at_0$

Path of light beam in accelerating frame

$t=0$ $t=t_0$ $t=2t_0$

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Is light bent by gravity?

- If we can't distinguish an accelerating reference frame from gravity...
- and light bends in an accelerating reference frame...
- then light must bend in a gravitational field

But light doesn't have any mass.
How can gravity affect light?

Maybe we are confused about what a straight line is

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Which of these is a straight line?

A. A

B. B

C. C

D. All of them

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Straight is shortest distance

- They are the shortest distances determined by wrapping string around a globe. On a globe, they are called 'great circles'. In general, geodesics.
- This can be a general definition of straight, and is in fact an intuitive one on curved surfaces
- It is the one Einstein used for the path of all objects in curved space-time
- The confusion comes in when you don't know you are on a curved surface.

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Mass and curvature

- General relativity says that any mass will give space-time a curvature
- Motion of objects in space-time is determined by that curvature
- Similar distortions to those we saw when we drew graphs in special relativity

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Idea behind geometric theory



- Matter bends space and time.
- Bending on a two-dimensional surface is characterized by the radius of curvature: r
- Einstein deduced that $1/r^2$ is proportional to the local energy and momentum density
- The proportionality constant is

$$\frac{8\pi G}{c^2}$$

- where G is Newton's constant

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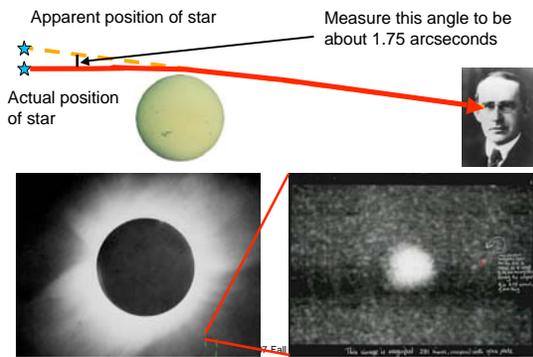
A test of General Relativity

- Can test to see if the path of light appears curved to us
- Local massive object is the sun
- Can observe apparent position of stars with and without the sun
- But need to block glare from sun

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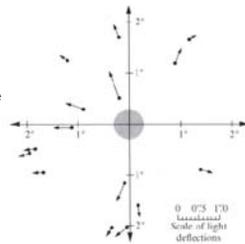
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Eddington and the Total Eclipse of 1919



Eddington's Eclipse Expedition 1919

- Eddington, British astronomer, went to Principe Island in the Gulf of Guinea to observe solar eclipse.
- After months of drought, it was pouring rain on the day of the eclipse
- Clouds parted just in time, they took photographic plates showing the location of stars near the sun.
- Analysis of the photographs back in the UK produced a deflection in agreement with the GR prediction

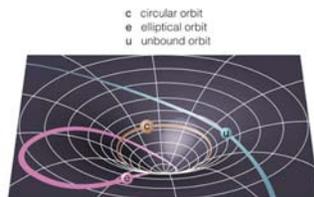


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Space is Curved

- Einstein said to picture gravity as a warp in space
- Kepler's Laws can all be explained by movement around these "puckers"
- Everything moving is affected, regardless of mass



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Other Consequences of GR

- Time dilation from gravity effects
- Gravitational Radiation!
 - Created when big gravity sources are moved around quickly
 - Similar to the electromagnetic waves that were caused by moving electron charges quickly
- Black Holes
- Expanding Universe (although Einstein missed the chance to predict it!)

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Gravitational time dilation

- Gravity warps both space and time!
- At 10,000 km above the Earth's surface, a clock should run 4.5 parts in 10^{10} faster than one on the Earth
- Comparing timing pulses from atomic oscillator clocks confirms the gravitational time dilation in 1976 to within 0.01%.
- Corrections are now standard in the synchronizing satellites
- This correction needed in addition to the special relativity correction for GPS

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Gravitational radiation

- When a mass is moved, the curvature of space-time changes
- If a mass is oscillated, ripples of space-time curvature carry the signal
- Gravitational radiation carries energy and momentum and wiggles mass in its path

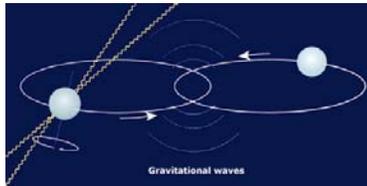


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Evidence for gravity waves

- In 1974, Joseph Taylor and his student Russell Hulse discovered a binary neutron star system losing energy as expected from gravitational radiation



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Direct detection of gravity waves

LIGO is a collection of large laser interferometers searching for gravity waves generated by exploding stars or colliding black holes



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The big bang

- In 1929 Observation of nearby and far away galaxies indicate that everything is receding from us.
 - Key physics needed to understand this is the simple Doppler shift of light waves. Waves from sources moving away from us are stretched out or lower frequency.
- Extrapolating backwards indicates that all the galaxies originated from the same source 14 billion years ago.
- In 1964 radiation from the early stages of that explosion was detected.
 - Again the Doppler shift was the key since the waves were shifted to low frequency - microwave

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This Years Nobel Prize

- For the universe to start small and expand space and time must be thing that can expand(or contract)
 - General relativity was key physics needed to understand that process
- However, a simple model of that would predict such a universe would not have clumps of matter(stars, galaxies)
- Unless those clumpings were present very early on
- This years Nobel prize was given to the people who designed the COBE experiment which was sensitive enough to see those clumpings in the CMB

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